

Nebencal code update

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Ubercal algorithm: relative calibration

AN IMPROVED PHOTOMETRIC CALIBRATION OF THE SLOAN DIGITAL SKY SURVEY IMAGING DATA

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$$m = m_{\text{ADU}} + a_{\alpha} - \left[k_{\beta} + \left(\frac{dk}{dt} \right)_{\beta} (t - t_{\beta,\text{ref}}) \right] x + f_{\gamma}(j), \quad (6)$$

$$\chi_i^2 = \sum_{j \in \mathcal{O}(i)} \left[\frac{m_i - m_{j,\text{ADU}} - a_{\alpha(j)} + k_{\beta(j)}(t)x - f_{\gamma(j)}}{\sigma_j} \right]^2, \quad (8)$$

$$\chi^2 = (\mathbf{A}\mathbf{p} - \mathbf{b})^t \mathbf{C}^{-1} (\mathbf{A}\mathbf{p} - \mathbf{b}), \quad (12)$$

$$\frac{d\chi^2}{d\mathbf{p}} = \mathbf{A}^t \mathbf{C}^{-1} \mathbf{A}\mathbf{p} - \mathbf{A}^t \mathbf{C}^{-1} \mathbf{b} = 0. \quad (14)$$

zeropoints
 dependence on position, airmass, etc.
 mag errors
 mag measurements

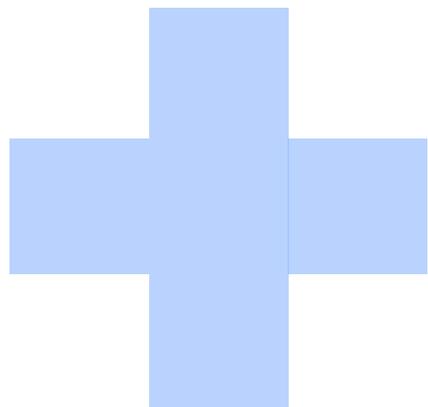
matrix equation:
 $\mathbf{A}\mathbf{x} = \mathbf{b}$

Algorithm

- Calibrate assuming that the weighted mean of “good” mags is the truth
 - Minimize **relative** offset between observations
- **Absolute** calibration
 - Include PreCam (and/or SDSS, ...) standards as if from one image
 - At the end, normalize all ZPs \rightarrow ZP(PreCam) to make a calibration pegged to the PreCam data

Nebencal algorithm

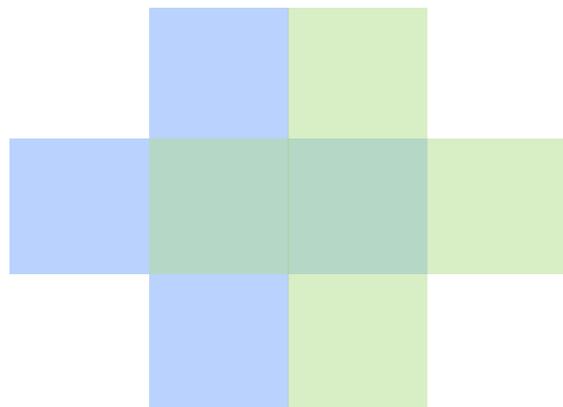
- Inspired by LSST's implementation
- Limit memory usage by pixelizing the sky via Healpix
- Calibrate each pixel plus neighbors in one übercal (~tens of degrees, customizable size)



- At the end, do another übercal to normalize these pixel regions together

Nebencal algorithm

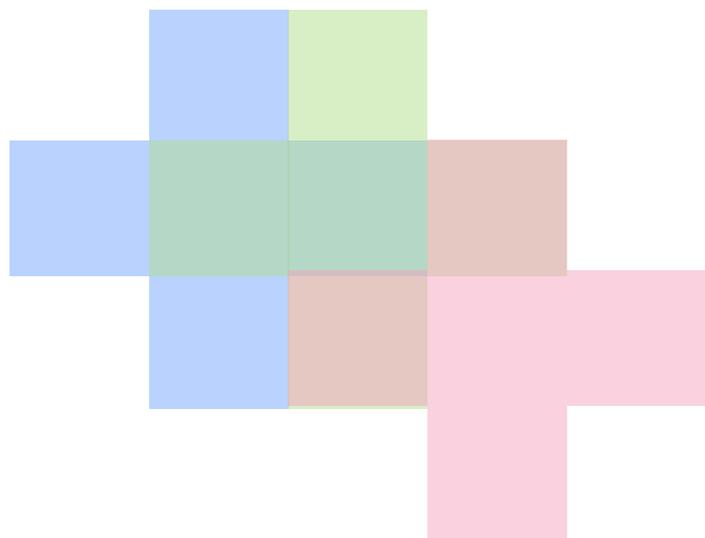
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Nebencal algorithm

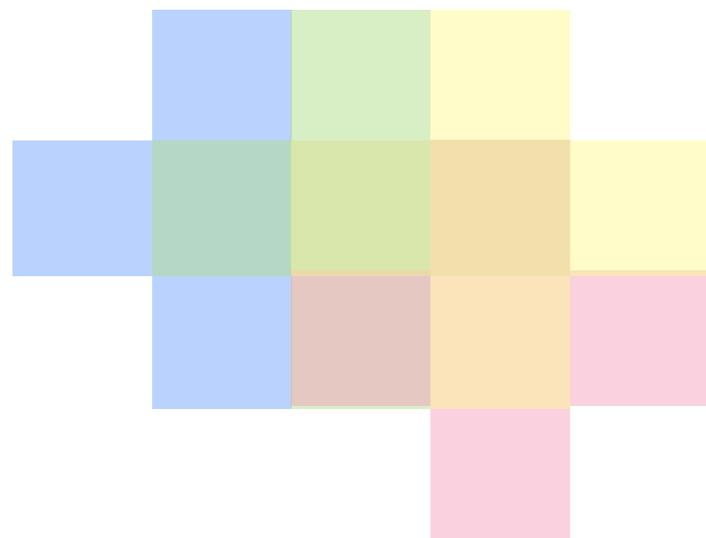
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Implementation: Config yaml file

general:

```
# what filter(s) do we want to calibrate?
filter : 'z'

# note that these input files depend on the filter!
precam_filename : "/Users/bauer/surveys/DES/precam/PreCamStandarStars/Z.Stand1percent.s"
zp_phot_filename : "/Users/bauer/surveys/DES/zp_phots/z.dat"

# where are the inputs?
globals_dir : '/Users/bauer/surveys/DES/y1p1/equatorial'
sdss_filename : "/Users/bauer/surveys/DES/y1p1/equatorial/sdss/SDSSDR10_SouthGalCap/stripe82_sample1.csv"

# include precam and/or sdss as standards?
use_precam : True
use_sdss : False

# what nside is used in the global object file pixelization scheme?
nside_file : 32
```

calibrations:

```
- id_string : 'exposureid'
  outfilename : 'nebencal_exp_zps_z'
  nside : 1
  max_dets : 2500
  require_standards : True

- id_string : 'image_id'
  outfilename : 'nebencal_img_zps_z'
  nside : 32
  max_dets : 500
  require_standards : False
```

Implementation: a couple of details

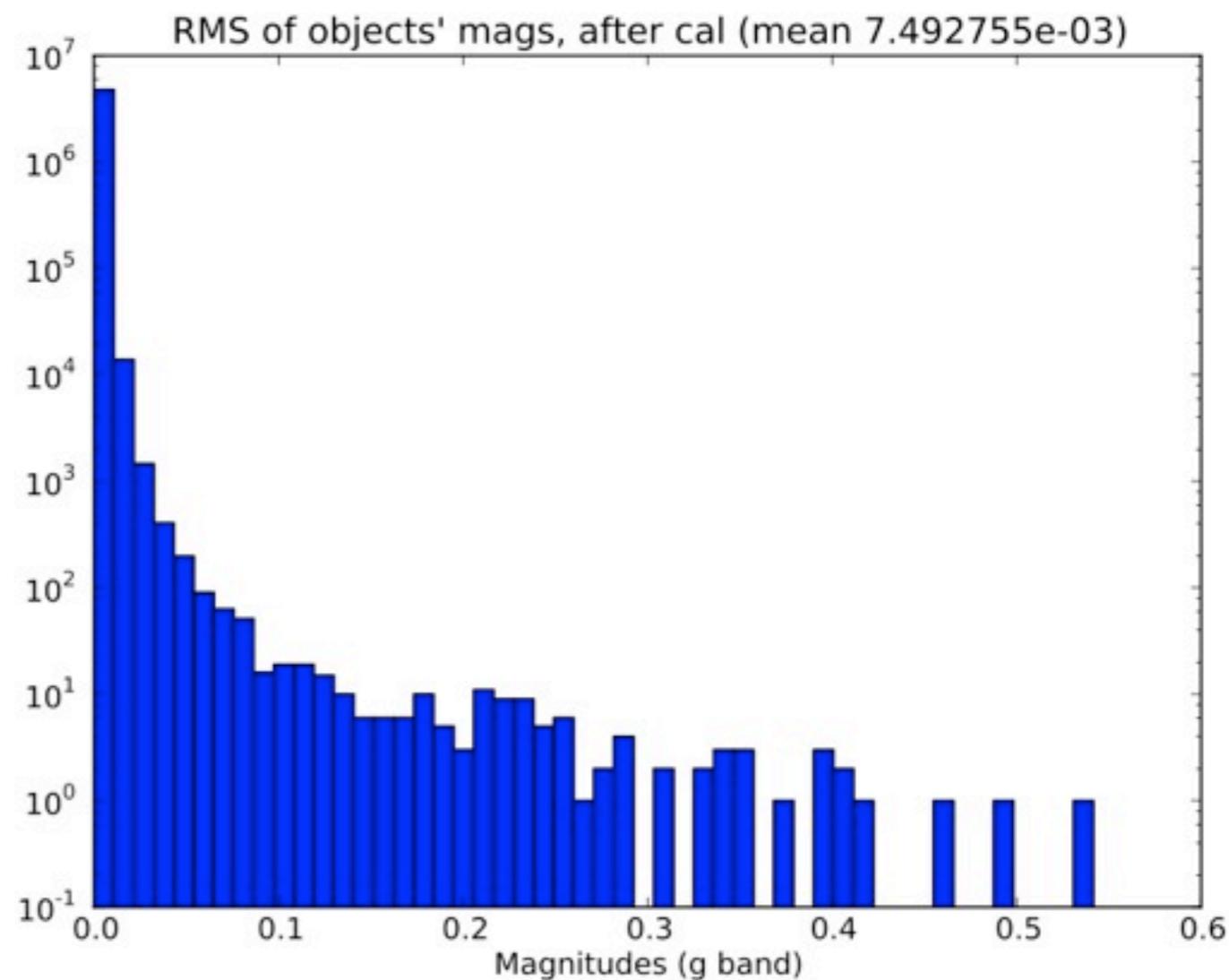
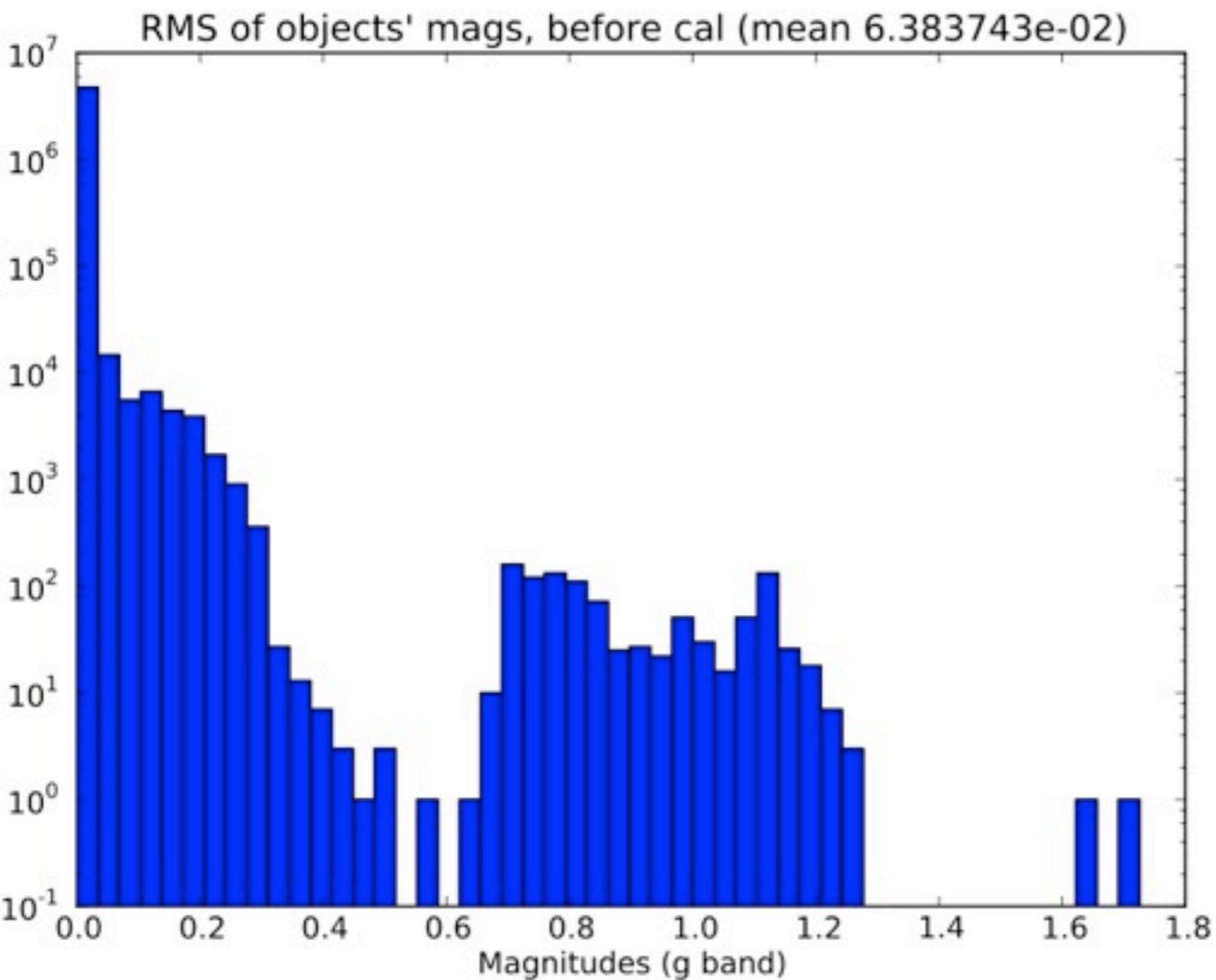
- First step, pre-calibration: match across exposures and make global objects. Saved in one binary file per Healpix pixel (in the highest pixel resolution I plan on using. The code degrades this file resolution to whatever is desired.)
 - Per detection: mag_psf & error, image & exposure ids, rasicam info. as yet unused: x, y, ccd, everything else in the DB...
- Next, just run: `nebencal.py config.yaml`
- I have my code in a git repository, let me know if you'd like access.

Implementation: a couple of details

- Written in python
- Uses sparse matrices to do the work: `scipy.sparse.linalg`
- The exposure calibration takes some (4-5ish) hours on one processor. The image calibration takes <1 hour. (For the equatorial region, per filter, without cuts on number of detections)
- Python multiprocessing module doesn't work with CBLAS functions. I may try harder to parallelize the code.
- Takes ~6GB memory per run (exposure calibration limited), but I think it should be 2-3x less...

g band RMS of each objects' mag measurements

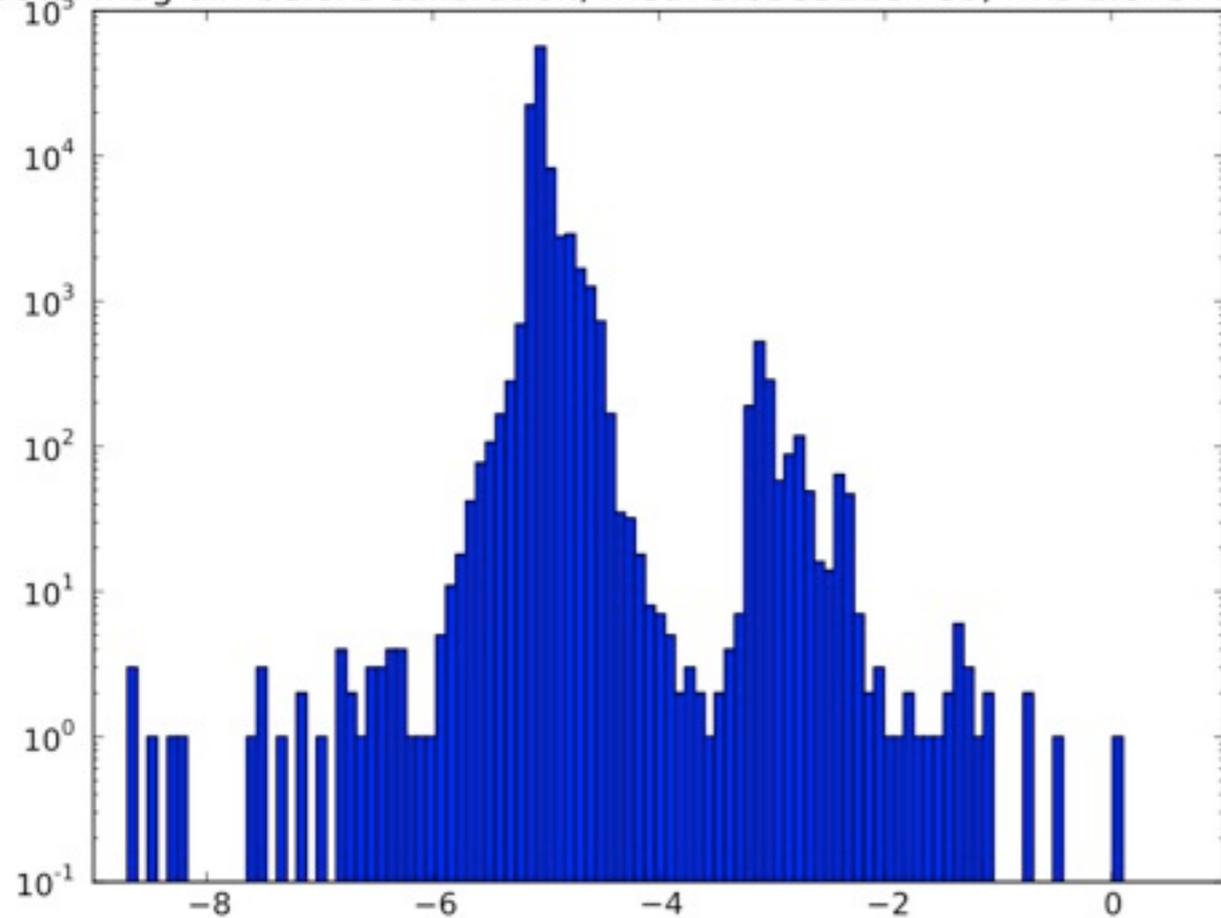
- Relative mag precision 6.4% calibrated to 0.75%



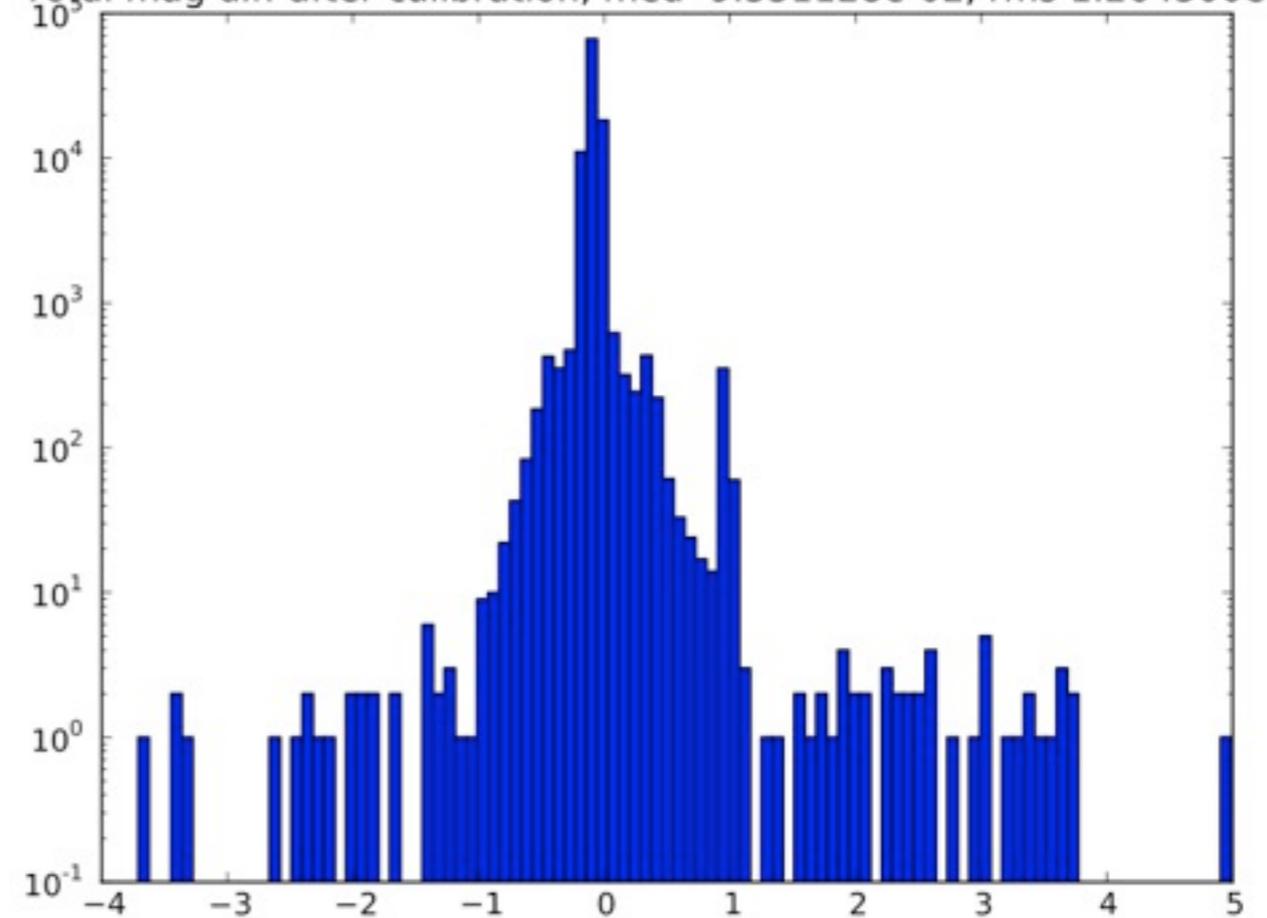
Absolute calibration is working (but sensitive to outliers)

- g-band using PreCam as standard objects, checking against SDSS as “truth”

Total mag diff before calibration, med $-5.058311e+00$, rms $2.873727e-01$



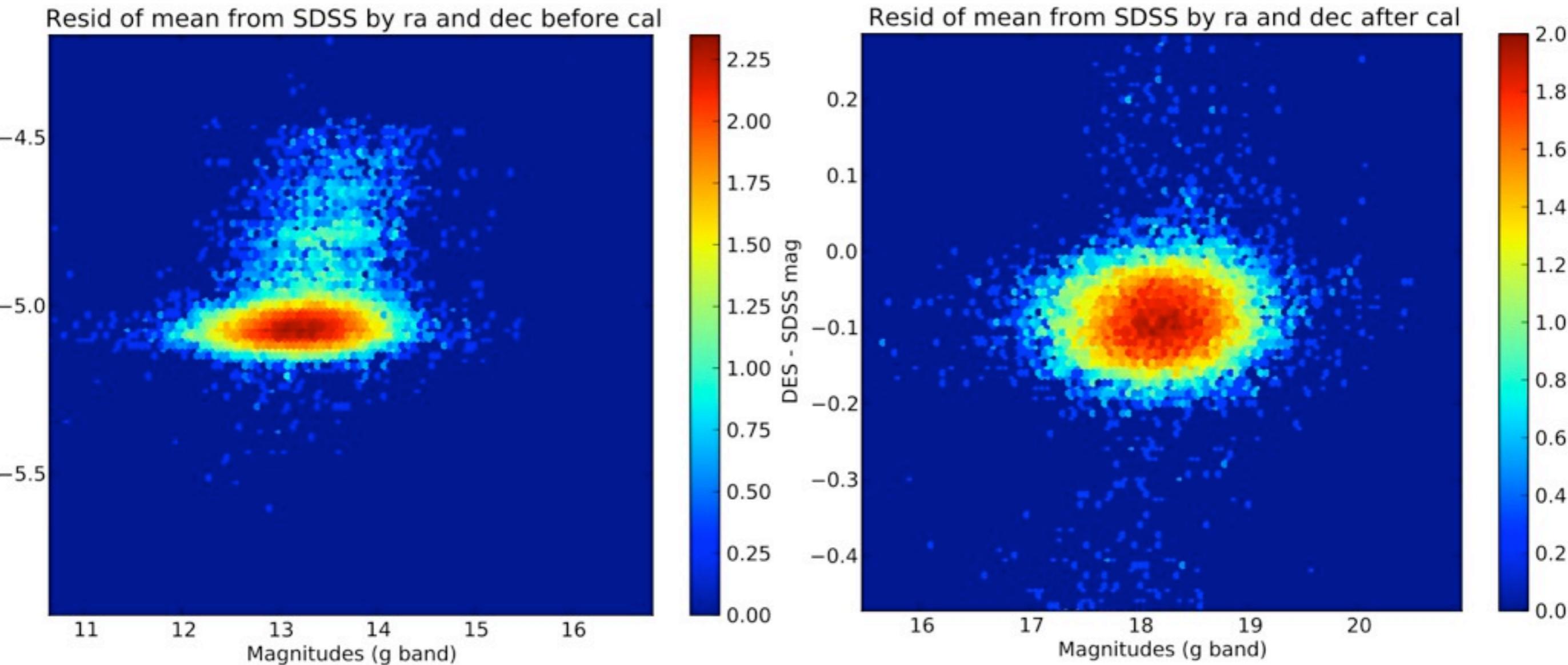
Total mag diff after calibration, med $-9.331128e-02$, rms $1.264300e-01$



- Note final median is still off due to outliers... a solution is in the works.

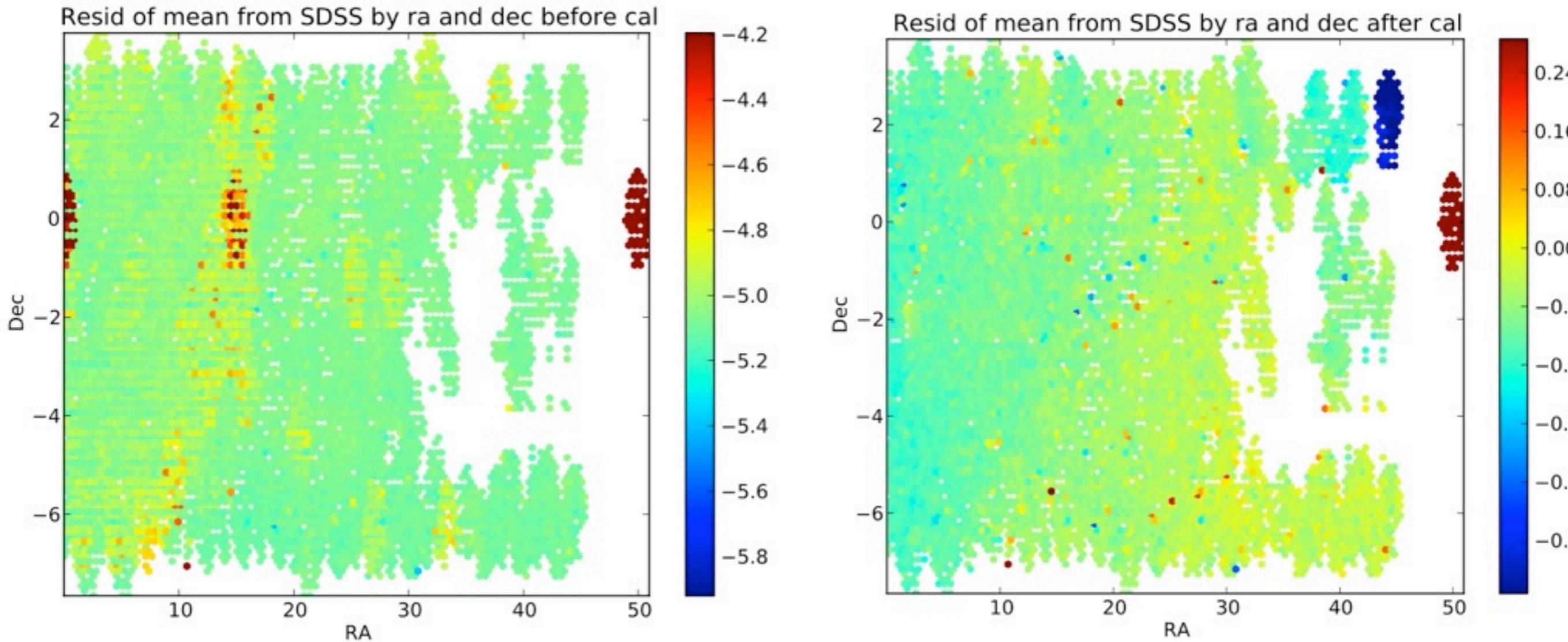
Absolute calibration vs DES magnitude

- DES-SDSS vs DES, before and after g-band calibration



Position dependence of absolute residuals

- The worst residuals are from disjoint regions

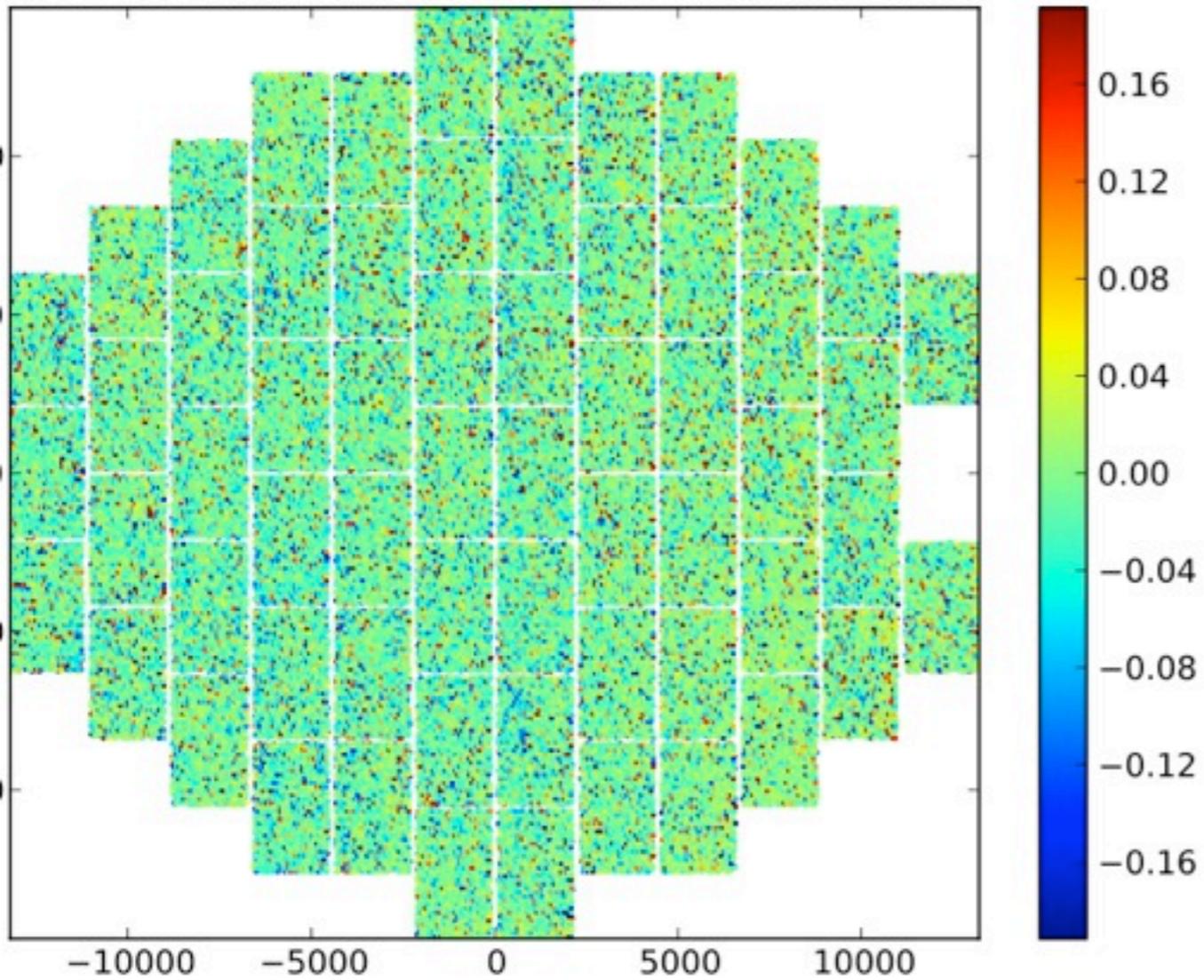


Plan to check for disjoint regions

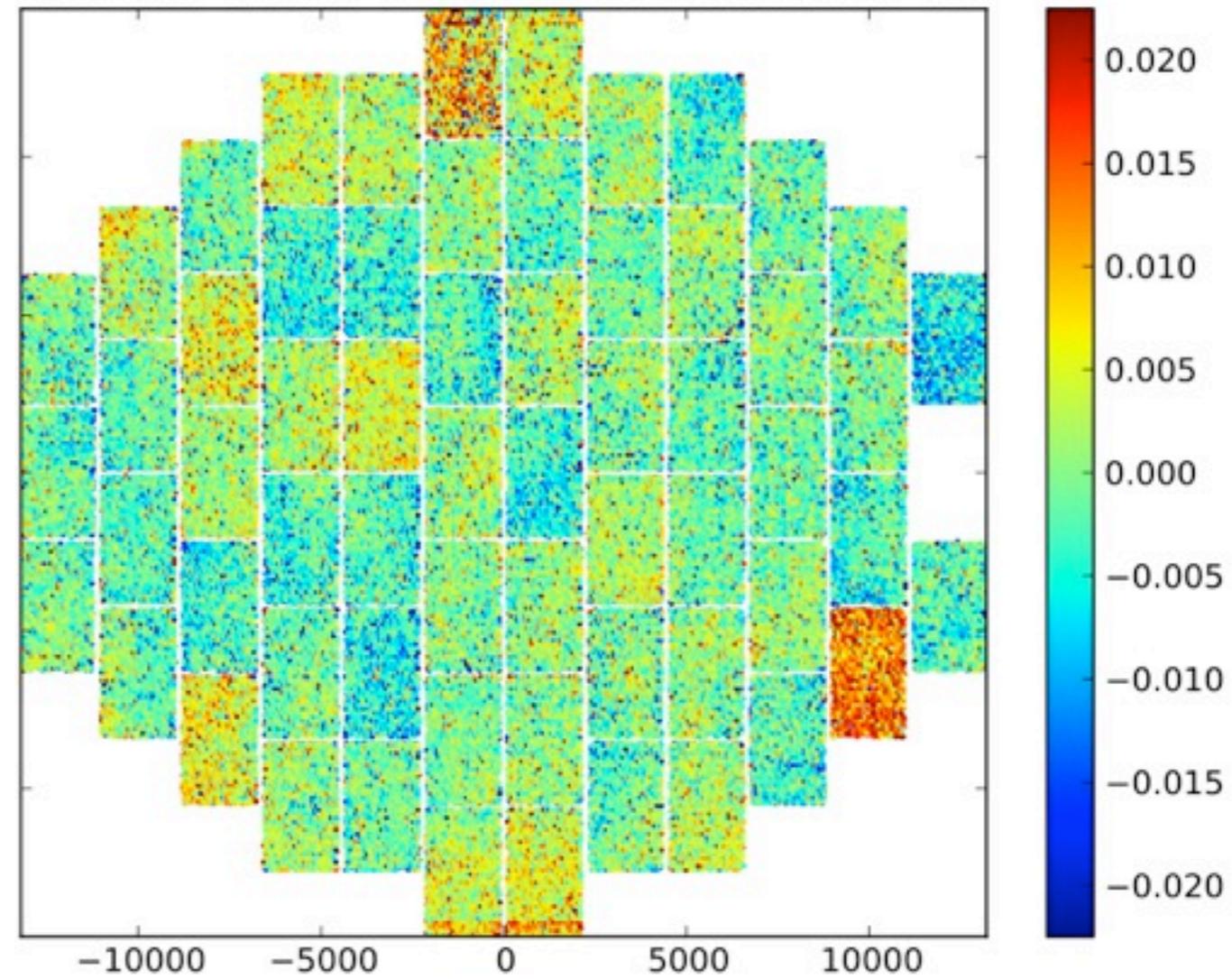
- I'm putting in a check to see if the graph of, e.g., image_id connections contains disconnected regions
 - `scipy.sparse.cs_graph_components`
- If we require coverage by standards and a region is disjoint from the standards, it will be removed from the calculation and flagged uncalibrated.
- Should help outlier rate and absolute calibration accuracy

Focal plane dependence of residuals: g band

Residual of indiv meas from DES mean before cal

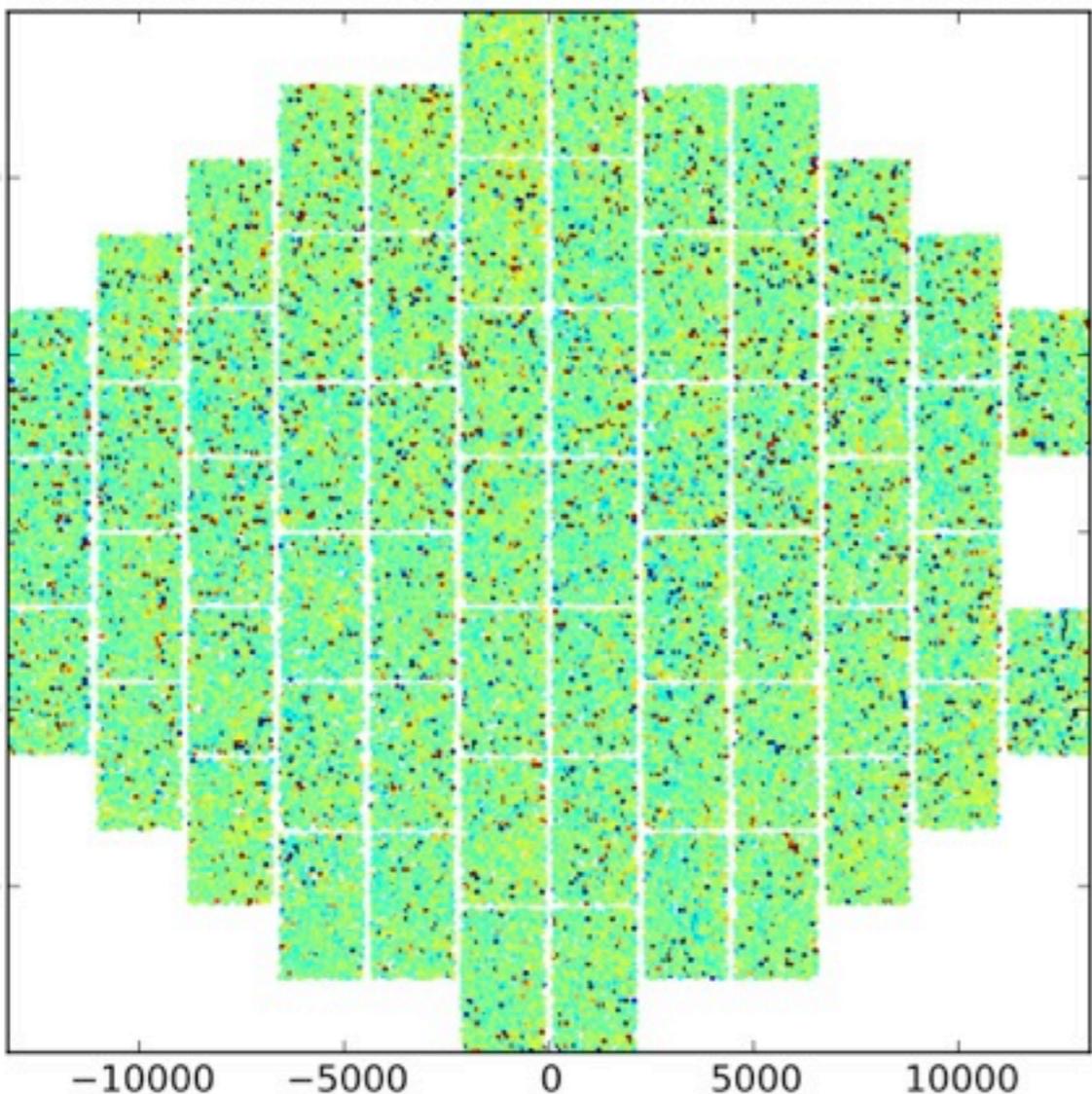


Residual of indiv meas from DES mean after cal

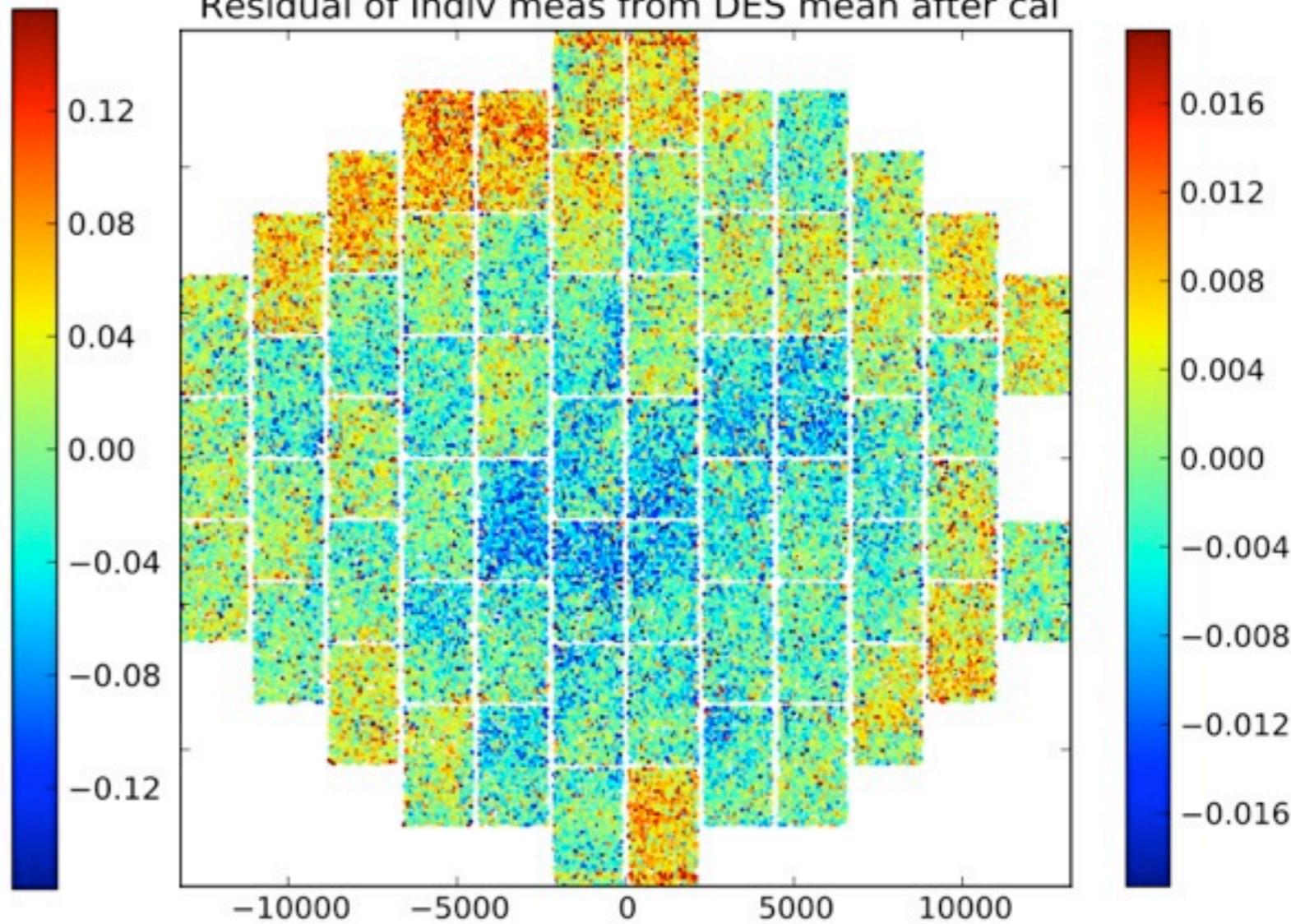


Focal plane dependence of residuals: i band

Residual of indiv meas from DES mean before cal

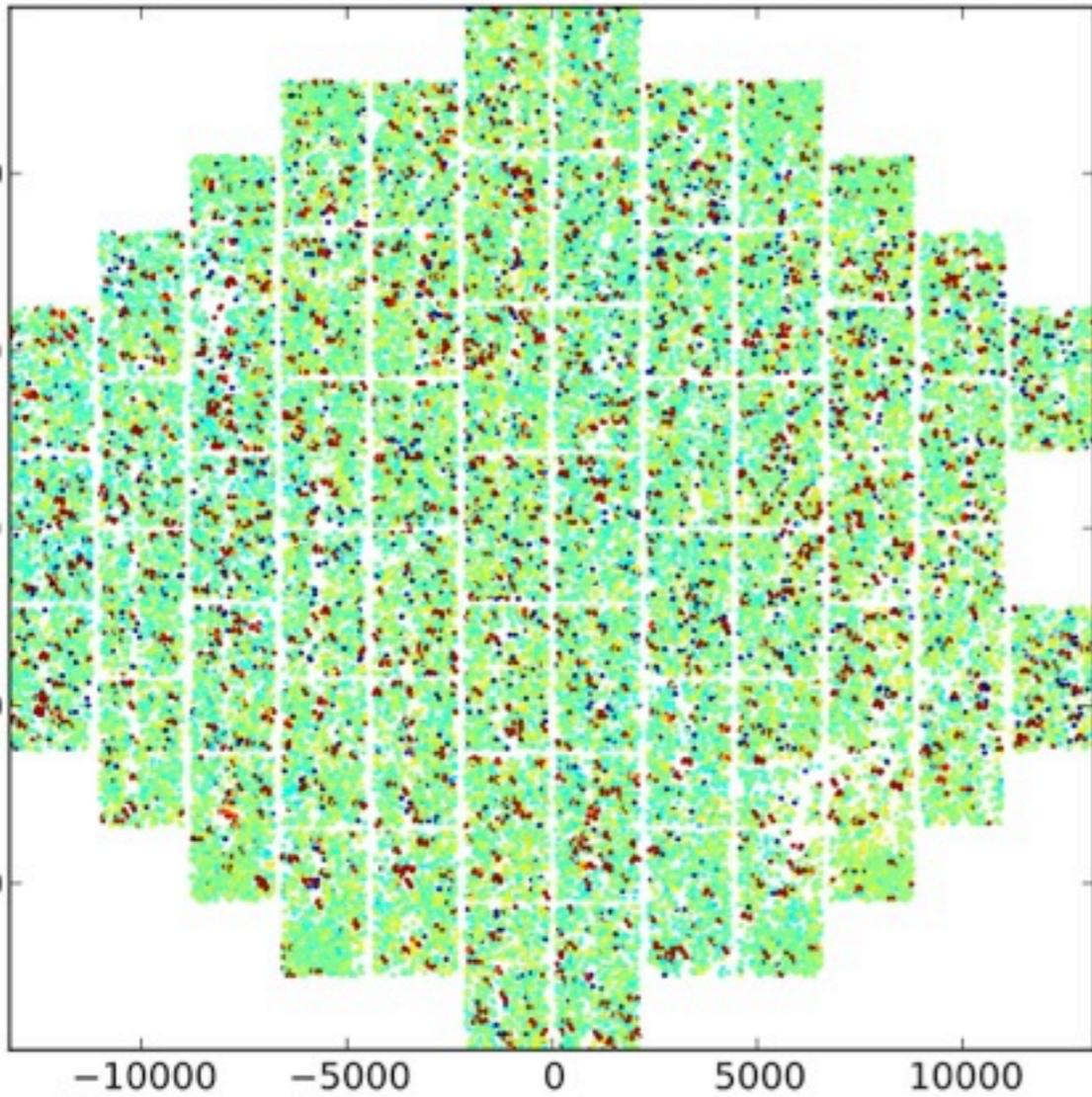


Residual of indiv meas from DES mean after cal

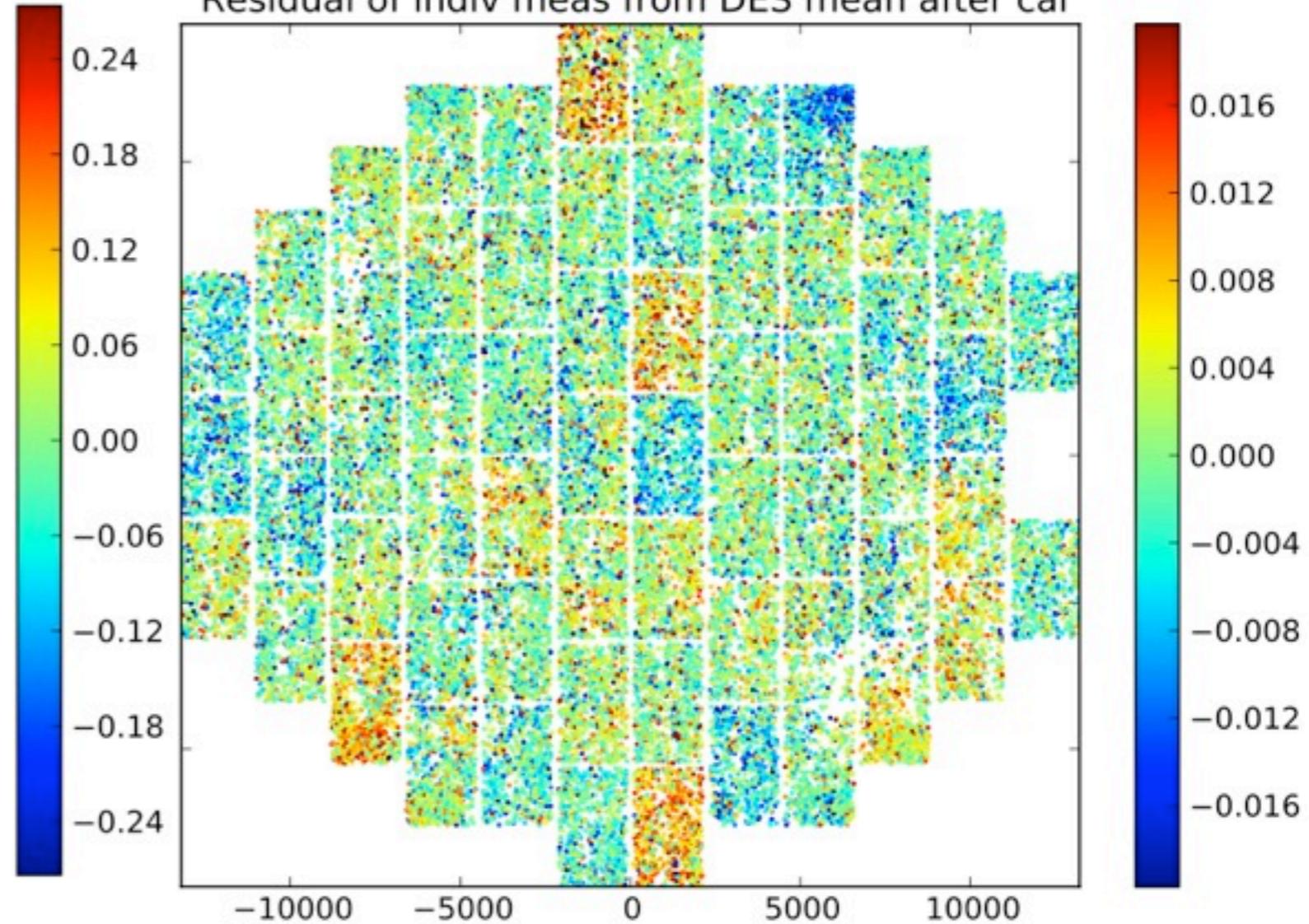


Focal plane dependence of residuals: g band

Residual of indiv meas from DES mean before cal



Residual of indiv meas from DES mean after cal



Future work

- outlier rejection on variable stars, bad measurements, etc.
- could compute zp error estimates
- exclude edges of chip? additional quality cuts?